



# MeO-ODISSEE/CESAR Optimize the injection into a fiber

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# The Meo-ODISSEE/CESAR project

### Objectives of the project CESAR

- Quantify and improve the injection into a single-mode optical fiber by controlling a tip-tilt mirror and by analyzing direct images
  - in the **visible** domain (R band and around),
  - in real-time (about 2 frames of delay in practice),
  - with partial AO correction (typically 8x8 actuators),
  - with **on-sky** conditions (by using the MeO/ODISSEE bench).

### Previous results with CESAR at CHARA

- TT correction is needed in the visible.
- Photocenter estimation on the closest bright speckle (not the brightest speckle).



# The CESAR module

### • The CESAR module

- One beam is directly imaged onto the detector.
- The other beam is injected into a single-mode fiber.
- Both images with and without the fiber are imaged on the same detector.
- A tip-tilt mirror (200Hz) enables to optimize the flux injected into the fiber.
- Spectral filters can select equal or different wavelengths for each beam.



# **TT mirror control & images analysis**

### Photocenter estimation

- **Reference pixel** (= **fiber location**) calibrated with an artificial source.
- **Photocenter** estimated for each short exposure image. 2 approaches:
- Photocenter of the **brightest speckle** on the whole image. The drawback is that the brightest speckle can be far from the reference pixel (due to a transition state), yielding to *speckle jump* and to TT correction unstability.
- 2. Photocenter of the closest bright speckle. On a small area (2-3 times the speckle size) centered on the reference pixel. This algorithm stabilize and maximizes the injection (by a factor 3). It relies on the AO efficiency.

### • Tip-tilt correction

- Command sent to the TT mirror to bring the speckle onto the reference pixel.
- TT\_correction = Gain \* Photocenter\_estimation
- Gain = 0.05 (should be as large as possible)
- Delay = 2.1 frames (should be as small as possible)

### • Performance evaluation

- **Coupling efficiency** into the fiber = 1.6 \* flux (fiber) / flux (direct)
- Different spectral filters for the direct beam and for the injected beam (to increase the throughput at the cost of a worse TT correction).

### **CESAR implementation on MeO-ODISSEE**



# **Timeline for Meo-ODISSEE/CESAR**

#### • February-March 2020 : Tests in laboratory

- Tests on a corner of the SPICA bench by using the collimated source.
- Tests in open loop (without the tip-tilt command, sent back to the supplier).
- Tests in close loop (with the tip-tilt command).

### • March 2020 : Implementation on MeO/ODISSEE

- Implementation of CESAR on an auxiliary arm in a corner of ODISSEE.
- Beam splitter 50/50 or 90/10 on the detection arm to split the light between:
  - the camera PCO sCMOS (used in operation for MeO-ODISSEE),
  - the camera ANDOR IXON 897 (used for CESAR).
- CESAR works independently from MeO/ODISSEE.

### • March-May 2020 : On-sky validation on MeO/ODISSEE

#### • May-June 2020 : Memorandum

- Achieved performance on Meo/ODISSEE (simulation + on-sky data).
- Expected performance on CHARA/SPICA (extrapolation by simulation).
- A step beyond : Control the TT mirror by analyzing the WFA data (???).

### **MeO-ODISSEE/CESAR**

# **The bench MeO/ODISSEE**

### • The MeO telescope

- 1.5-m aperture
- High-contrast NGS &
- Wide-field LGS mode

### • The bench ODISSEE

 It corrects in real time the wavefront distortions caused by the atmospheric turbulence.



- It mainly consists of an **adaptive optics** and an **imaging system** :
- Tip-tilt mirror : correction of modes related tilting of the wave (87% in the dynamics of corrections);
- Deformable Mirror (88 actuators) : correction of the remaining modes (13% in the dynamics of corrections);
- Wave Front Analyzer : Microlenses array (8x8) and high-speed camera (< 1500Hz) with a CCD of 240x240 pixels (24µm)</li>
- **Detection** : Field iris diaphragm of 1" and sCMOS photodetector.