

Solar System Science Before and After Gaia

# Shape models and densities of asteroids

THE POST-GAIA ERA

Benoît Carry

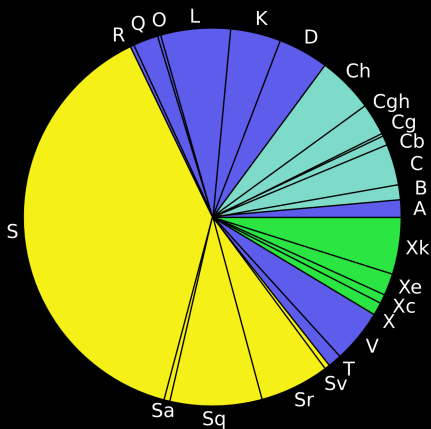
European Space Agency

with

S. Mouret, J. Āurech, M. Kaasalainen

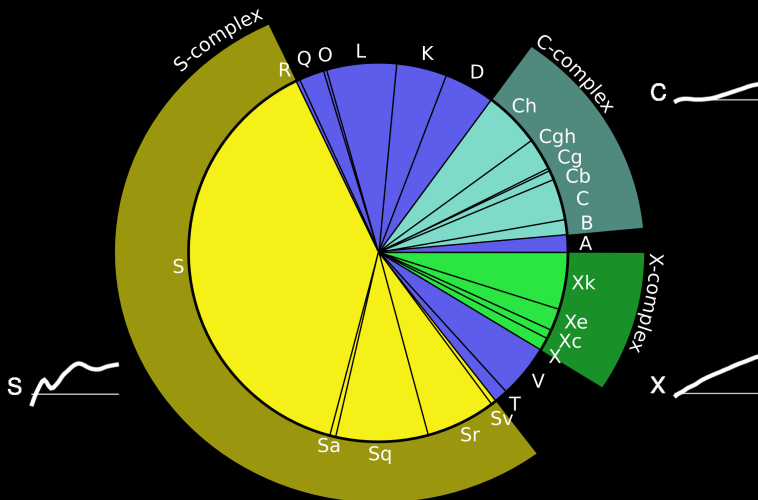


# Density & Composition



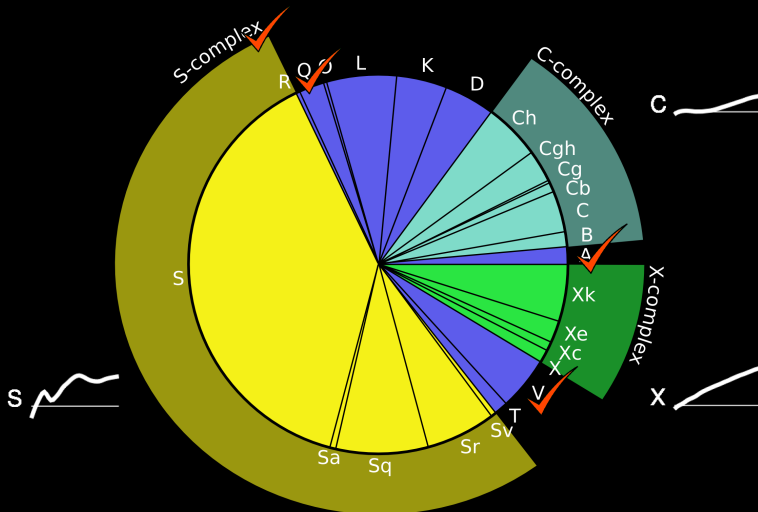
based on DeMeo et al. 2009

# Density & Composition



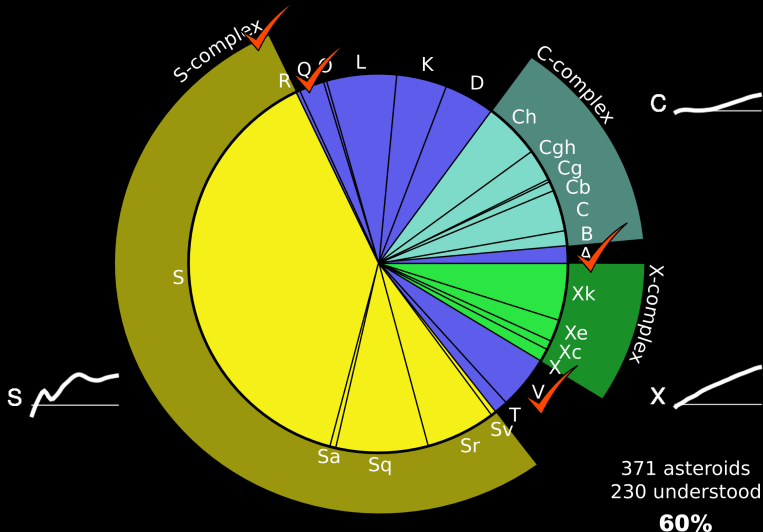
based on DeMeo et al. 2009

# Density & Composition



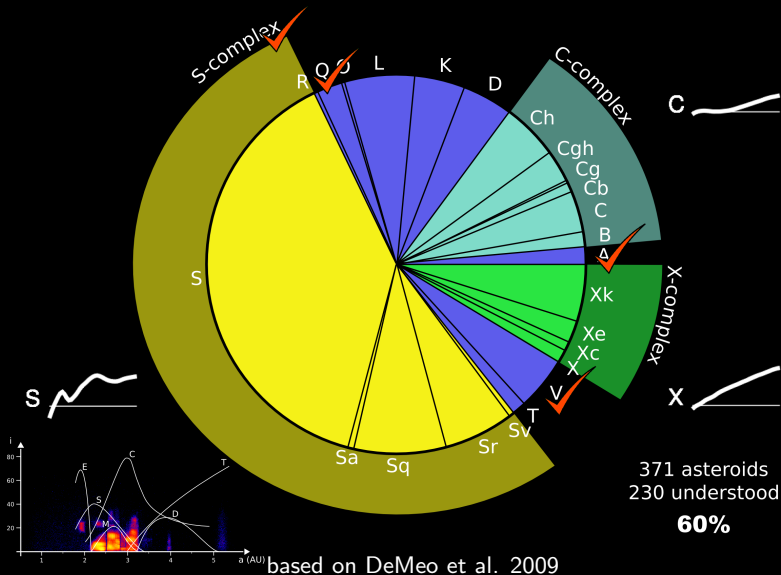
based on DeMeo et al. 2009

# Density & Composition

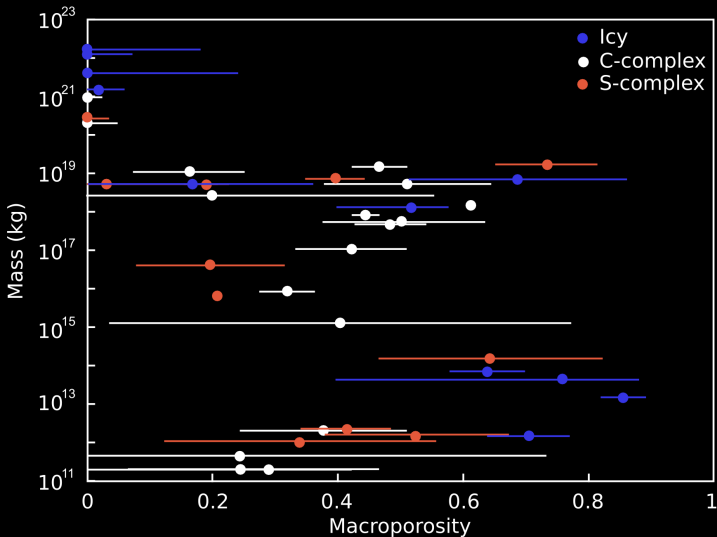


based on DeMeo et al. 2009

# Density & Composition

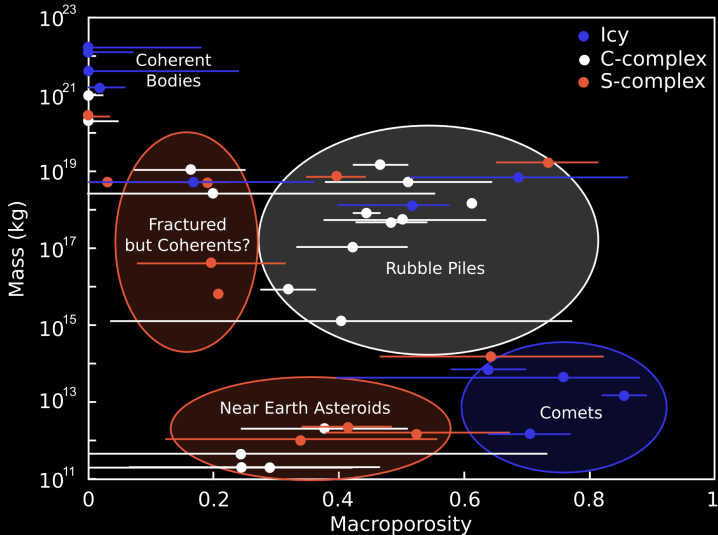


# Density & Structure



adapted from Britt et al. 2006

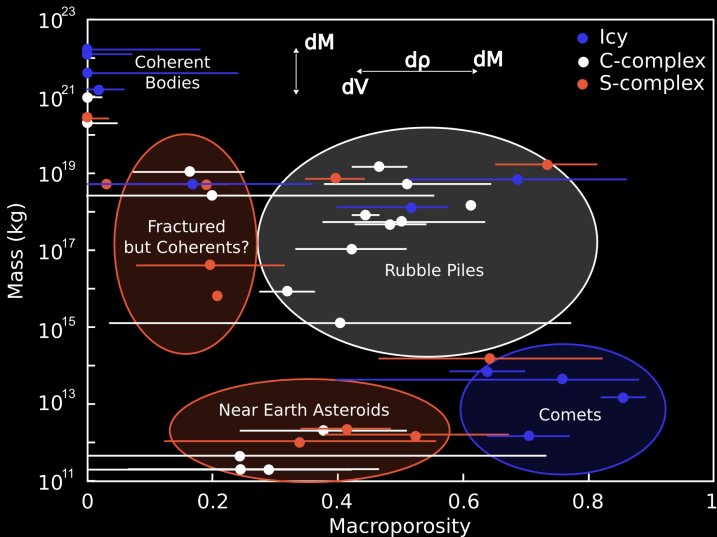
# Density & Structure



adapted from Britt et al. 2006



# Density & Structure



adapted from Britt et al. 2006

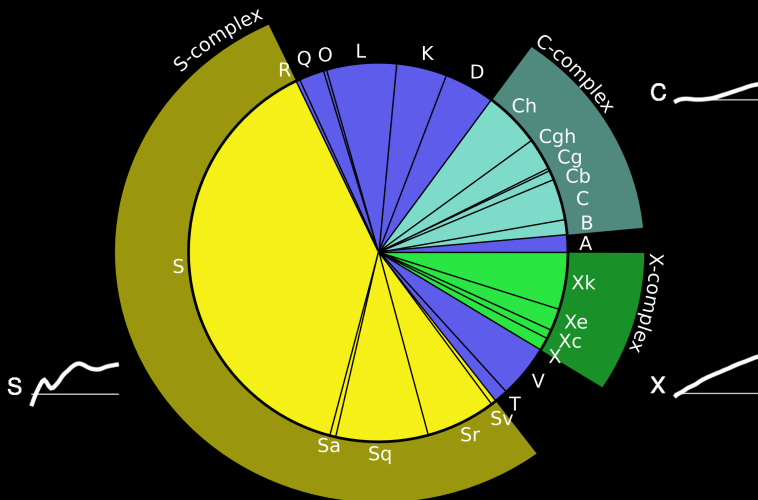
# Density & Troubles

$$\rho = \frac{M}{V}$$

# Mass measurements

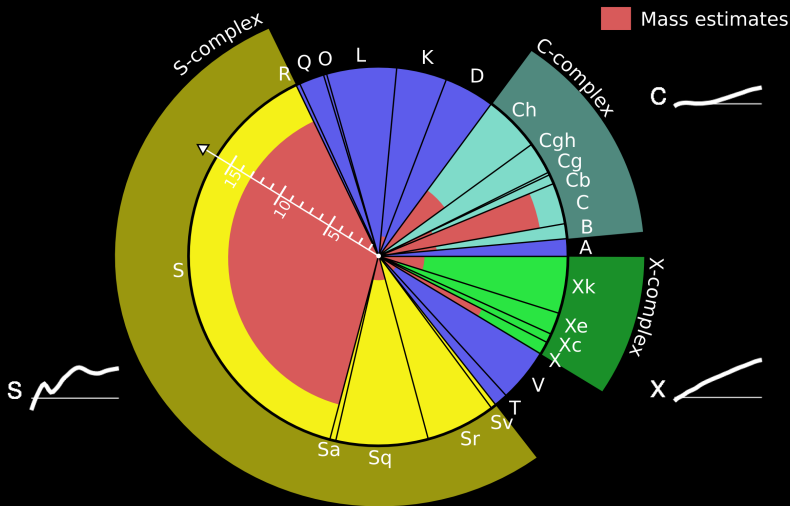
$$\rho = \frac{M}{V}$$

# Mass & Taxonomy



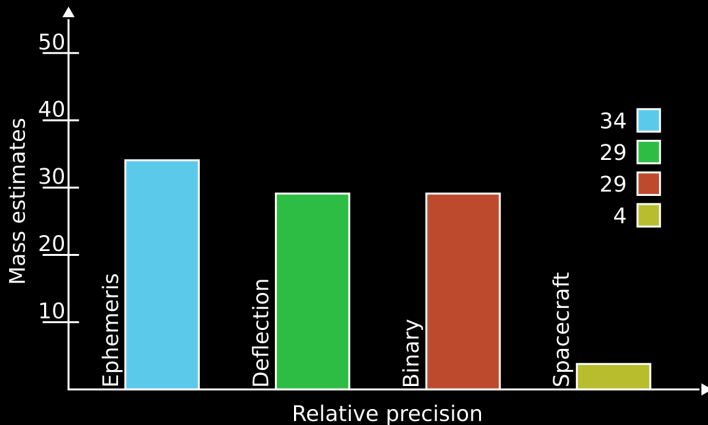
based on DeMeo et al. 2009

# Mass & Taxonomy

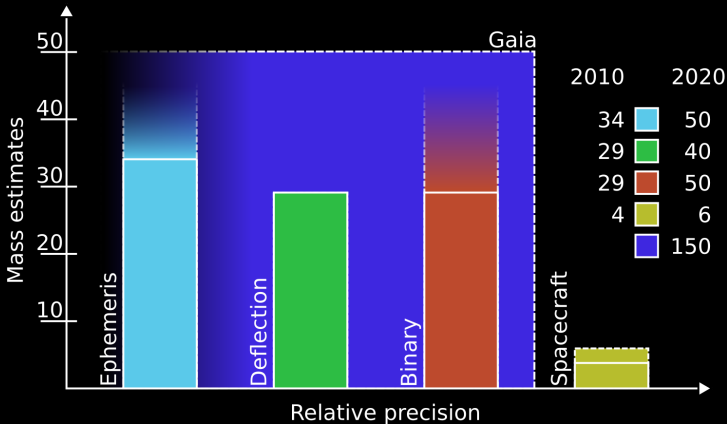


based on DeMeo et al. 2009 & Baer et al. 2011

# Mass & Uncertainty



# Mass & Gaia



# Mass & Gaia

- Breakthrough in mass estimates
- **Deflection & Binary**
- About **200-300** masses
- Coherent set
- Span taxonomy



# Volume estimates

$$\rho = \frac{M}{V}$$

# Volume & Troubles

$$\rho = \frac{\mathcal{M}}{\mathcal{V}}$$

- **100-300** masses
- **2228** diameters IRAS
- **10 000** diameters Spitzer
- **100 000** diameters WISE
- **10 000** diameters Gaia

## Volume & Troubles

$$\rho = \frac{\mathcal{M}}{\mathcal{V}}$$

$$\begin{aligned} \frac{d\rho}{\rho} &= \frac{d\mathcal{M}}{\mathcal{M}} + \frac{d\mathcal{V}}{\mathcal{V}} \\ &= \frac{d\mathcal{M}}{\mathcal{M}} + 3\frac{d\mathcal{R}}{\mathcal{R}} \end{aligned}$$

- **100-300** masses
- **2228** diameters IRAS
- **10 000** diameters Spitzer
- **100 000** diameters WISE
- **10 000** diameters Gaia

## Volume & Troubles

$$\rho = \frac{\mathcal{M}}{\mathcal{V}}$$

$$\begin{aligned} \frac{d\rho}{\rho} &= \frac{d\mathcal{M}}{\mathcal{M}} + \frac{d\mathcal{V}}{\mathcal{V}} \\ &= \frac{d\mathcal{M}}{\mathcal{M}} + 3\frac{d\mathcal{R}}{\mathcal{R}} \end{aligned}$$

- **100-300** masses
- **2228** diameters IRAS
- **10 000** diameters Spitzer
- **100 000** diameters WISE
- **10 000** diameters Gaia

**Size**  
is the  
**Limiting**  
**Factor**

# Accurate Volumes

## 1. Direct measurements

- **WYSIWYG**  $\neq$  model-dependant
- ▶ Disk-resolved imaging
- ▶ Stellar occultations

## 2. Realistic 3-D shapes

## 3. Completeness of geometries

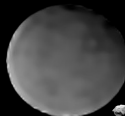
(21) Lutetia



(135) Hertha



(1) Ceres



1 Euro



@ 2km

Tanga &amp; Delbò 2007

Carry et al. 2008

Durech et al. 2011

# Accurate Volumes

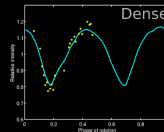
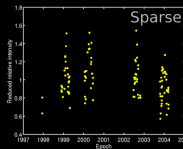
## 1. Direct measurements

- **WYSIWYG**  $\neq$  model-dependant
- ▶ Disk-resolved imaging
- ▶ Stellar occultations

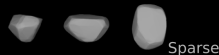
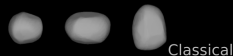
## 2. Realistic 3-D shapes

- Assumptions  $\Rightarrow$  **bias**
- Concavity  $\leftrightarrow$  Volume
- ▶ Lightcurves (dense & sparse)

## 3. Completeness of geometries



Models



Kaasalainen & Ďurech  
Hanuš et al. 2011

$\Rightarrow$  Cellino, Hanuš, Carbognani

# Accurate Volumes

## 1. Direct measurements

- **WYSIWYG**  $\neq$  model-dependant
- ▶ Disk-resolved imaging
- ▶ Stellar occultations

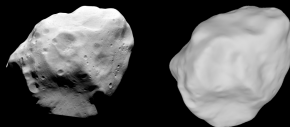
## 2. Realistic 3-D shapes

- Assumptions  $\Rightarrow$  **bias**
- Concavity  $\leftrightarrow$  Volume
- ▶ Lightcurves (dense & sparse)

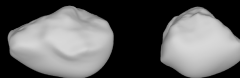
## 3. Completeness of geometries

- **Extensive** approach
- ▶ Thermal radiometry
- ▶ Interferometry
- ▶ Radar echo

(21) Lutetia



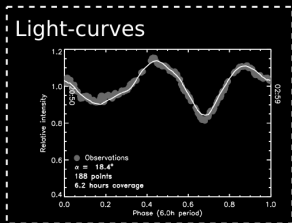
GREAT!



Not so great....

Tanga et al. 2003  
Busch et al. 2009  
Delbò et al. 2009

# Volume with KOALA

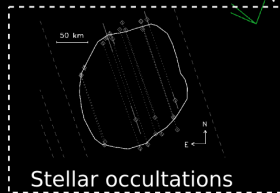


Indirect measurement  
Shape  
Spin

Direct measurement  
Size  
Shape

Knitted  
Occultation  
Adaptive optics  
Light-curve  
Analysis

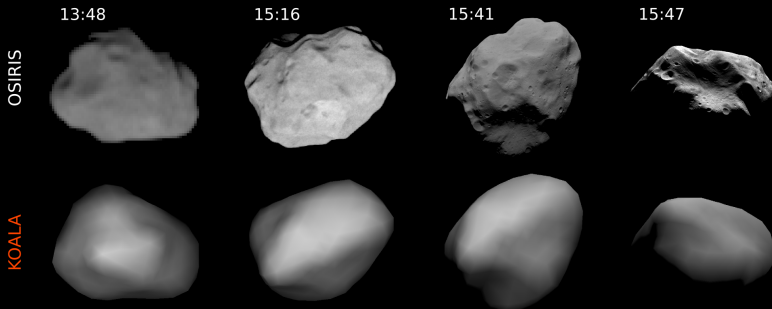
Direct measurement  
Size  
Shape



Kaasalainen, IPI 2011



# Volume with KOALA



Pre-Flyby model  
**KOALA**

Carry et al. 2010

vs. Rosetta  
Shape: **2 km**

Carry et al. 2011

Accuracy  
**Volume: 15%**

# KOALA Development

- **KOALA 1.0:** Kaasalainen, Ďurech & Carry
  - ▶ **Lightcurves** (dense & sparse)
  - ▶ Disk-resolved **imaging**
  - ▶ Stellar **occultations**
  - ▶ Code being cleaned for release
- **KOALA 1.x:** Current development
  - ▶ **Interferometry** (optical): *FGS on HST, JWST*
  - ▶ **Thermal** radiometry: *IRAS, Spitzer, WISE*
  - ▶ Kaasalainen, Ďurech, Delbò, Carry
- **KOALA 2.0:** Planned development
  - ▶ **Interferometry** (thermal): *VLTI, LBT*
  - ▶ **Radar:** *Arecibo, LOFAR+EISCAT3d*

# Gaia's Contribution

$$\rho = \frac{M}{V}$$

# Gaia's Contribution

- **Ultra-high precision astrometry**

- ▶ Mass determination: **Deflection** & Binary
- ▶ **150** estimates – spanning taxonomy
- ▶ Ground-based support Colas
- ▶ Ephemeris (e.g. INPOP) Fienga
- ▶ **N** extra estimates?

- **Very-high precision photometry**

- ▶ 3-D shape reconstruction Cellino, Hanuš, Carbognani
- ▶ Minor contribution for volume

- ▶ **Ground-based support**

- ▶ **Lightcurves** (dense) campaigns
- ▶ Dedicated **imaging** on large telescopes
- ▶ Stellar **occultations** Tanga, Delbò